# On a new type of solving procedure for Euler-Poisson equations (rigid body rotation over the fixed point) 

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#### Abstract

In this paper, we proceed to develop a new approach which was formulated first in Ershkov (Acta Mech 228(7):2719-2723, 2017) for solving Poisson equations: a new type of the solving procedure for EulerPoisson equations (rigid body rotation over the fixed point) is suggested in the current research. Meanwhile, the Euler-Poisson system of equations has been successfully explored for the existence of analytical solutions. As the main result, a new ansatz is suggested for solving Euler-Poisson equations: the Euler-Poisson equations are reduced to a system of three nonlinear ordinary differential equations of first order in regard to three functions $\Omega_{i}(i=1,2,3)$; the proper elegant approximate solution has been obtained as a set of quasi-periodic cycles via re-inversing the proper elliptical integral. So the system of Euler-Poisson equations is proved to have analytical solutions (in quadratures) only in classical simplifying cases: (1) Lagrange's case, or (2) Kovalevskaya's case or (3) Euler's case or other well-known but particular cases.


## 1 Introduction, equations of motion

Euler-Poisson equations, describing the dynamics of rigid body rotation, are known to be one of the famous problems in classical mechanics.

In accordance with [1-3], Euler equations describe the rotation of a rigid body in a frame of reference fixed in the rotating body for the case of rotation over the fixed point as below (at given initial conditions):

$$
\left\{\begin{array}{l}
I_{1} \frac{\mathrm{~d} \Omega_{1}}{\mathrm{~d} t}+\left(I_{3}-I_{2}\right) \cdot \Omega_{2} \cdot \Omega_{3}=P\left(\gamma_{2} c_{0}-\gamma_{3} b_{0}\right),  \tag{1.1}\\
I_{2} \frac{\mathrm{~d} \Omega_{2}}{\mathrm{t} t}+\left(I_{1}-I_{3}\right) \cdot \Omega_{3} \cdot \Omega_{1}=P\left(\gamma_{3} a_{0}-\gamma_{1} c_{0}\right), \\
I_{3} \frac{\mathrm{~d} \Omega_{3}}{\mathrm{~d} t}+\left(I_{2}-I_{1}\right) \cdot \Omega_{1} \cdot \Omega_{2}=P\left(\gamma_{1} b_{0}-\gamma_{2} a_{0}\right),
\end{array}\right.
$$

where $I_{i} \neq 0$ are the principal moments of inertia ( $i=1,2,3$ ) and $\Omega_{i}$ are the components of the angular velocity vector along the proper principal axis; $\gamma_{i}$ are the components of the weight of mass $P$ and $a_{0}, b_{0}, c_{0}$ are the appropriate coordinates of the center of masses in a frame of reference fixed in the rotating body (in regard to the absolute system of coordinates $X, Y, Z)$.

[^0]The Poisson equations for the components of the weight in a frame of reference fixed in the rotating body (in regard to the absolute system of coordinates $X, Y, Z$ ) can be presented as below [4,5]:

$$
\left\{\begin{array}{l}
\frac{\mathrm{d} \gamma_{1}}{\mathrm{~d} t}=\Omega_{3} \gamma_{2}-\Omega_{2} \gamma_{3}  \tag{1.2}\\
\frac{\mathrm{~d} \gamma_{2}}{\mathrm{~d} t}=\Omega_{1} \gamma_{3}-\Omega_{3} \gamma_{1} \\
\frac{\mathrm{~d} \gamma_{3}}{\mathrm{~d} t}=\Omega_{2} \gamma_{1}-\Omega_{1} \gamma_{2}
\end{array}\right.
$$

besides, we should present the invariants (first integrals of motion) as

$$
\left\{\begin{array}{l}
\gamma_{1}^{2}+\gamma_{2}^{2}+\gamma_{3}^{2}=1  \tag{1.3}\\
I_{1} \cdot \Omega_{1} \cdot \gamma_{1}+I_{2} \cdot \Omega_{2} \cdot \gamma_{2}+I_{3} \cdot \Omega_{3} \cdot \gamma_{3}=\text { const. }=C_{0} \\
\frac{1}{2}\left(I_{1} \cdot \Omega_{1}^{2}+I_{2} \cdot \Omega_{2}^{2}+I_{3} \cdot \Omega_{3}^{2}\right)+P\left(a_{0} \gamma_{1}+b_{0} \gamma_{2}+c_{0} \gamma_{3}\right)=\text { const. }=C_{1}
\end{array}\right.
$$

## 2 Derivation of the invariants (first integrals) of motion

According to [6], let us recall how to derive the invariants (1.3). From (1.1), (1.2) we obtain

$$
\begin{aligned}
& \frac{\frac{\mathrm{d}}{\mathrm{~d} t}\left(I_{1} \frac{\Omega_{1}^{2}}{2}+I_{2} \frac{\Omega_{2}^{2}}{2}+I_{3} \frac{\Omega_{3}^{2}}{2}\right)}{P}=\Omega_{1} \cdot\left(\gamma_{2} c_{0}-\gamma_{3} b_{0}\right)+\Omega_{2} \cdot\left(\gamma_{3} a_{0}-\gamma_{1} c_{0}\right)+\Omega_{3} \cdot\left(\gamma_{1} b_{0}-\gamma_{2} a_{0}\right), \\
& \Rightarrow \quad-\frac{\frac{\mathrm{d}}{\mathrm{~d} t}\left(I_{1} \frac{\Omega_{1}^{2}}{2}+I_{2} \frac{\Omega_{2}^{2}}{2}+I_{3} \frac{\Omega_{3}^{2}}{2}\right)}{P}=a_{0} \cdot\left(\Omega_{3} \gamma_{2}-\Omega_{2} \gamma_{3}\right)+b_{0} \cdot\left(\Omega_{1} \gamma_{3}-\Omega_{3} \gamma_{1}\right)+c_{0} \cdot\left(\Omega_{2} \gamma_{1}-\Omega_{1} \gamma_{2}\right), \\
& \Rightarrow \quad \frac{\mathrm{d}}{\mathrm{~d} t}\left(I_{1} \frac{\Omega_{1}^{2}}{2}+I_{2} \frac{\Omega_{2}^{2}}{2}+I_{3} \frac{\Omega_{3}^{2}}{2}\right)+a_{0} \cdot P \cdot \frac{\mathrm{~d} \gamma_{1}}{\mathrm{~d} t}+b_{0} \cdot P \cdot \frac{\mathrm{~d} \gamma_{2}}{\mathrm{~d} t}+c_{0} \cdot P \cdot \frac{\mathrm{~d} \gamma_{3}}{\mathrm{~d} t}=0 .
\end{aligned}
$$

So we have obtained the third integral of (1.3) in [7]. For obtaining the second integral of (1.3) in [7], we multiply each of Eqs. (1.1) by $\gamma_{i}$ accordingly, but also each of equations of (1.2) by ( $I_{i} \cdot \Omega_{i}$ ) accordingly; then we add as below:

$$
\left\{\begin{array}{l}
\left(I_{1} \cdot \gamma_{1} \cdot \frac{\mathrm{~d} \Omega_{1}}{\mathrm{~d} t}+\gamma_{1} \cdot\left(I_{3}-I_{2}\right) \cdot \Omega_{2} \cdot \Omega_{3}\right)+\left(I_{1} \cdot \Omega_{1} \frac{\mathrm{~d} \gamma_{1}}{\mathrm{~d} t}\right)=\gamma_{1} \cdot P\left(\gamma_{2} c_{0}-\gamma_{3} b_{0}\right)+I_{1} \cdot \Omega_{1} \cdot\left(\Omega_{3} \gamma_{2}-\Omega_{2} \gamma_{3}\right) \\
\left(I_{2} \cdot \gamma_{2} \cdot \frac{\mathrm{~d} \Omega_{2}}{\mathrm{~d} t}+\gamma_{2} \cdot\left(I_{1}-I_{3}\right) \cdot \Omega_{3} \cdot \Omega_{1}\right)+\left(I_{2} \cdot \Omega_{2} \frac{\mathrm{~d} \gamma_{2}}{\mathrm{~d} t}\right)=\gamma_{2} \cdot P\left(\gamma_{3} a_{0}-\gamma_{1} c_{0}\right)+I_{2} \cdot \Omega_{2} \cdot\left(\Omega_{1} \gamma_{3}-\Omega_{3} \gamma_{1}\right) \\
\left(I_{3} \cdot \gamma_{3} \cdot \frac{\mathrm{~d} \Omega_{3}}{\mathrm{~d} t}+\gamma_{3} \cdot\left(I_{2}-I_{1}\right) \cdot \Omega_{1} \cdot \Omega_{2}\right)+\left(I_{3} \cdot \Omega_{3} \frac{\mathrm{~d} \gamma_{3}}{\mathrm{~d} t}\right)=\gamma_{3} \cdot P\left(\gamma_{1} b_{0}-\gamma_{2} a_{0}\right)+I_{3} \cdot \Omega_{3} \cdot\left(\Omega_{2} \gamma_{1}-\Omega_{1} \gamma_{2}\right)
\end{array}\right.
$$

Having done this, we add all the three equations above:

$$
I_{1} \cdot \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\Omega_{1} \cdot \gamma_{1}\right)+I_{2} \cdot \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\Omega_{2} \cdot \gamma_{2}\right)+I_{3} \cdot \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\Omega_{3} \cdot \gamma_{3}\right)=0 .
$$

So we have obtained the second integral of (1.3) in [7].
The first integral of (1.3) is trivial, but belongs to Poisson equations only: to obtain it, we multiply each of equations of (1.2) on $\gamma_{i}$ accordingly, then add them (the constant of integration is chosen equal to 1 , due to trigonometric sense of the presenting solution in absolute system of coordinates via Euler angles):

$$
\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\gamma_{1}^{2}\right)+\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\gamma_{2}^{2}\right)+\frac{1}{2} \frac{\mathrm{~d}}{\mathrm{~d} t}\left(\gamma_{3}^{2}\right)=0
$$

As we can see, 2 of 3 proper additional invariants above are obtained by using of all the 6 EP-equations (including Poisson equations).

But aforesaid argument is not sufficient for solving the EP-equations: indeed, the system of equations (1.1)-(1.2) is supposed to not be equivalent to the system of equations (1.1) along with all the invariants (1.3) (Dr. Hamad H. Yehya, personal communications) for some particular cases, as was suggested earlier in [7,8]. The rather complex case, which describes the motion of the constrained rigid body around a fixed point, was considered in the comprehensive article [9].

So, for solving the system of equations (1.1)-(1.2), we first solve the Poisson equations (1.2).


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